Attitude Stabilization for Magnetically Actuated Spacecraft using Rotation Matrices

Fabio Celani School of Aerospace Engineering







- quaternion feedback (Q-F) and the unwinding phenomenon
- rotation matrix feedback (RM-F)

• comparison between Q-F and RM-F for a Cubesat case study

Quaternion and Attitude Rate Feedback



quaternion $(\mathbf{q}_{\mathbf{v}}, q_4)$ $\mathbf{q}_{\mathbf{v}} = [q_1 \ q_2 \ q_3]^T$ vector part q_4 scalar part

body frame = inertial frame \Leftrightarrow either $(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, 1)$ or $(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, -1)$

for fully-actuated spacecraft $\mathbf{T} = -k_p \mathbf{q}_{\mathbf{v}} - k_d \boldsymbol{\omega}$ (PD-Quat) $\Downarrow (\mathbf{T}_{disturbance} = \mathbf{0})$ $(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, 1)$ almost globally asymptotically stable $(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, -1)$ unstable

Unwinding Phenomenon



Chaturvedi, Sanyal, McClamroch. Rigid-body attitude control. Control Systems Magazine. 2011

Attitude Stabilization for Magnetically Actuated Spacecraft using Rotation Matrices Fabio Celani

Rotation Matrix and Attitude Rate Feedback



rotation matrix ${\bf R}$



for fully-actuated spacecraft

$$\mathbf{T} = -\frac{k_p}{4} \sum_{i=1}^{3} (\mathbf{e_i} \times \mathbf{R}^T \mathbf{e_i}) - k_d \boldsymbol{\omega} \qquad [\mathbf{e_1} \ \mathbf{e_2} \ \mathbf{e_3}] = \mathbf{I}_{3 \times 3} \qquad (\mathsf{PD-RM})$$
$$\Downarrow \ (\mathbf{T}_{disturbance} = \mathbf{0})$$

$\mathbf{R} = \mathbf{I}_{3\times 3}$ almost globally asymptotically stable

Chaturvedi, Sanyal, McClamroch. Rigid-body attitude control. Control Systems Magazine. 2011





Page 7

Tigrisat



 $J_x = J_y = 4.09 \cdot 10^{-2} \text{ kg m}^2$ $J_z = 6.5 \cdot 10^{-3} \text{ kg m}^2$ circular orbit altitude = 629 km $T_{orbit} = 5832 \text{ sec}$ inclination = 97° RAAN = 68.5° camera pointing along z_b 3 orthogonal magnetorquers

Back to Cubesat World



objective: stabilize attitude so that body frame is aligned with orbital frame (camera pointing to Earth)

Quaternion and Attitude Rate Feedback



B geomagnetic field at spacecraft $\mathbf{m}_{coils} = -\mathbf{B} \times (K_p \mathbf{q_v} + K_d \boldsymbol{\omega}_{bo})$ (MPD-Quat) $\Downarrow (\mathbf{T}_{disturbance} = \mathbf{0})$ quaternion $(\mathbf{q}_{\mathbf{v}}, q_4)$

vector part $\mathbf{q}_{\mathbf{v}} = [q_1 \ q_2 \ q_3]^T$ scalar part q_4

body frame = orbital frame

either $(q_v, q_4) = (0, 1)$ or $(q_v, q_4) = (0, -1)$



 $(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, 1)$ locally asymptotically stable

 $(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, -1)$ unstable \Rightarrow unwinding phenomenon

Rotation Matrix and Attitude Rate Feedback



rotation matrix ${\bf R}$

body frame = orbital frame

$$\mathbf{R} = \mathbf{I}_{3 \times 3}$$

 $\begin{array}{l} \mathbf{GeoCentric Inertial (GCI) frame} \\ \mathbf{B} \ \ \mathbf{geomagnetic field at spacecraft} \\ \mathbf{m}_{coils} = -\mathbf{B} \times \left[\frac{K_p}{4} \sum_{i=1}^{3} (\mathbf{e_i} \times \mathbf{R}^T \mathbf{e_i}) + K_d \boldsymbol{\omega} \right] \\ \mathbf{e_1} \ \mathbf{e_2} \ \mathbf{e_3} \right] = \mathbf{I}_{3 \times 3} \\ \mathbf{M}_{coils} = \mathbf{I}_{3 \times 3} \\ \begin{array}{l} \mathbf{M}_{coils} = -\mathbf{B} \times \left[\frac{K_p}{4} \sum_{i=1}^{3} (\mathbf{e_i} \times \mathbf{R}^T \mathbf{e_i}) + K_d \boldsymbol{\omega} \right] \\ \mathbf{M}_{coils} = -\mathbf{B} \times \left[\frac{K_p}{4} \sum_{i=1}^{3} (\mathbf{e_i} \times \mathbf{R}^T \mathbf{e_i}) + K_d \boldsymbol{\omega} \right] \\ \mathbf{M}_{coils} = -\mathbf{B} \times \left[\frac{K_p}{4} \sum_{i=1}^{3} (\mathbf{e_i} \times \mathbf{R}^T \mathbf{e_i}) + K_d \boldsymbol{\omega} \right] \\ \mathbf{M}_{coils} = \mathbf{M}_{coils} \begin{bmatrix} 293.4863 & 0.5515 & -9.7049 \\ -0.0069 & 299.8118 & -4.1120 \\ 4.8505 & -0.1118 & 299.8613 \end{bmatrix} \\ \mathbf{M}_{coils} = \mathbf{M}_{coils} \begin{bmatrix} 1.8 \cdot 10^4 & 0 & 0 \\ 0 & 1.8 \cdot 10^4 & 0 \\ 0 & 0 & 1.8 \cdot 10^4 \end{bmatrix} \\ \mathbf{M}_{coils} = \mathbf{M}_{coils} \begin{bmatrix} 1.8 \cdot 10^4 & 0 & 0 \\ 0 & 0 & 1.8 \cdot 10^4 \end{bmatrix} \\ \mathbf{M}_{coils} = \mathbf{M}_{coils} \begin{bmatrix} \mathbf{M}_{coils} + \mathbf{M}_{coils} \\ \mathbf{M}_{coils} = \mathbf{M}_{coils} \end{bmatrix}$

 $\mathbf{R} = \mathbf{I}_{3 \times 3}$ locally **asymptotically stable** (no unwinding phenomenon)

 $\mathbf{T}_{disturbance}
eq \mathbf{0}$



Attitude Stabilization for Magnetically Actuated Spacecraft using Rotation Matrices

Fabio Celani

 $\mathbf{T}_{disturbance} \neq \mathbf{0}$



Attitude Stabilization for Magnetically Actuated Spacecraft using Rotation Matrices Fabio Celani

Monte Carlo campaign

1000 simulation runs random initial attitude random $\omega(0)$ with $\|\omega(0)\| \le 20 \text{ deg/sec}$

results

mean ITAE MPD-Quat = **2.56 10¹¹ deg sec²** mean ITAE MPD-RM = **2.34 10¹¹ deg sec² (-9%)** number of runs in which (ITAE MPD-RM < ITAE MPD-Quat) = **49%**

mean "energy" consumption MPD-Quat = **9.94** A² m² sec mean "energy" consumption MPD-RM = **8.71** A² m² sec (-12%) number of runs in which ("energy" MPD-RM < "energy" MPD-Quat) = **96%**

Conclusion

 comparison between MPD-Quat and MPD-RM attitude control laws for a CubeSat

 Monte Carlo campaign shows that the two control laws are comparable in terms of speed of convergence

 Monte Carlo campaign shows that MPD-RM leads to lower "energy" consumption

linearizations

 $\mathbf{m}_{coils} = -\mathbf{B} \times (K_p \mathbf{q}_{\mathbf{v}} + K_d \boldsymbol{\omega}_{bo})$ (MPD-Quat)

linearization about
$$(\mathbf{q}_{\mathbf{v}}, q_4) = (\mathbf{0}, 1)$$

$$\mathbf{m}_{coils} = -\mathbf{B} \times \begin{bmatrix} \frac{K_p}{2} \boldsymbol{\zeta} + K_d \boldsymbol{\omega} \end{bmatrix} \quad \begin{array}{l} \boldsymbol{\zeta} = [\phi \ \theta \ \psi]^T \\ \mathbf{3}\text{-2-1 Euler angles} \end{array}$$

initial condition close to desired attitude



Page 15